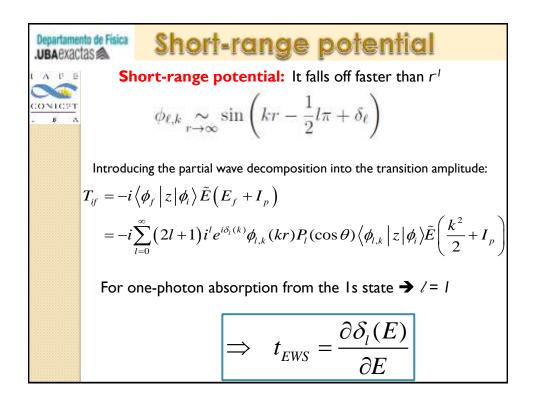
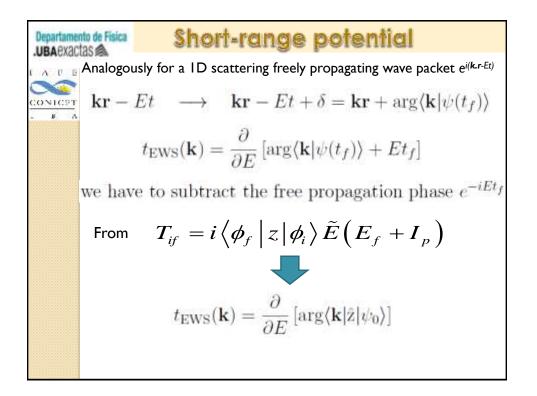


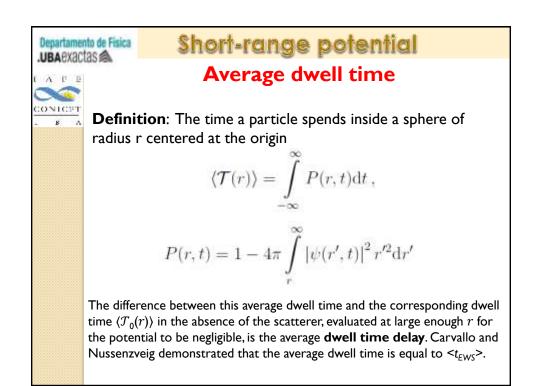
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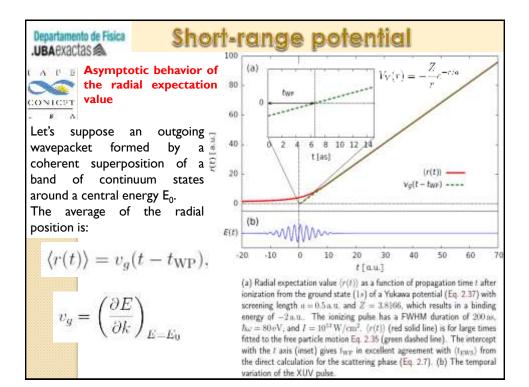


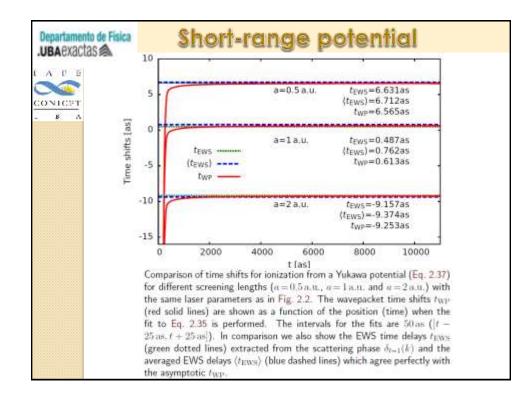
EXAMPLE 1 Short-range potential Non-perturbative analysis: $i \frac{\partial}{\partial t} |\psi(t)\rangle = \hat{H} |\psi(t)\rangle$ $\hat{H} = \frac{\hat{p}^2}{2} + V_{sr}(r) + \hat{H}_{el}$ Asymptotically, $H = p^2/2$, with solutions $\langle \mathbf{r} | \mathbf{k} \rangle = e^{i\mathbf{k}\mathbf{r}}$ $\langle \mathbf{r} | \psi(t_f) \rangle = \langle \mathbf{r} | \int d^3k | \mathbf{k} \rangle \langle \mathbf{k} | \psi(t_f) \rangle$ $= \int d^3k e^{i(\mathbf{k}\mathbf{r} + \arg(\langle \mathbf{k} | \psi(t_f) \rangle))} |\langle \mathbf{k} | \psi(t_f) \rangle|$



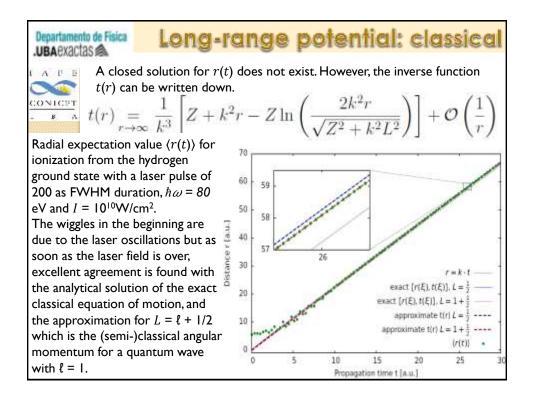
EXAMPLE Short-range potential W. Brenig and R. Haag. Allgemeine Quantentheorie der Stoßprozesse Fortschr. Phys. 7, 183 (1959). C. A. A. de Carvalho and H. M. Nussenzveig. Time delay. Physics Reports **364**, 83 (2002). proved that the expectation value of the position of a wave packet $\langle r \rangle = \langle v \rangle t + c + O(t^{-1}), \quad t \to \infty$ $\langle v \rangle = \frac{1}{m} \int \left| \tilde{\psi}(\hat{\mathbf{k}}) \right|^2 \left| \hat{\mathbf{k}} \right| d^3k = \frac{\langle k \rangle}{m}$ $c = -\hbar \int \left| \tilde{\psi}(\hat{\mathbf{k}}) \right|^2 \frac{\partial}{\partial k} \arg \tilde{\psi}(\hat{\mathbf{k}}) d^3k$ $= -\hbar \left\langle \frac{\partial}{\partial k} \arg \tilde{\psi}(\hat{\mathbf{k}}) \right\rangle = -\hbar \left\langle v \frac{\partial}{\partial E} \arg \tilde{\psi}(\hat{\mathbf{k}}) \right\rangle$ $t_{\rm BH} = \hbar \left\langle \frac{\partial}{\partial E} \arg \tilde{\psi}(\hat{\mathbf{k}}) \right\rangle = \langle t_{\rm EWS} \rangle$



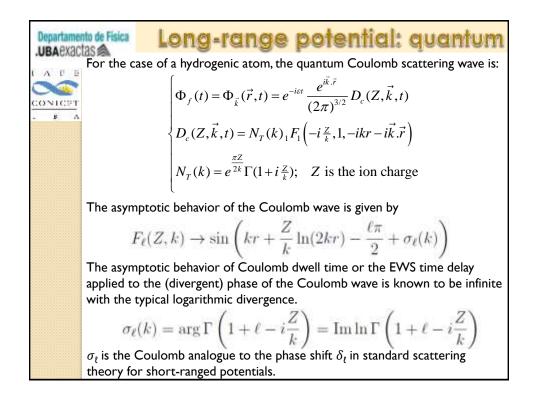




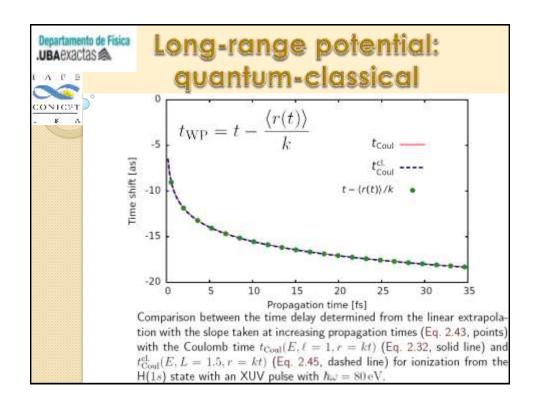
Departamento de Física Long-range potential: classical .UBAexactas So far, we have considered short range potentials. 0 B Let's consider now a classical Coulomb (Kepler) trajectory. CONTEPT The effective potential energy is given by в à. $U_{\rm eff}(r) = -\frac{Z}{r} + \frac{L^2}{2r^2}$ For an energy E < 0 the particle has a finite motion in an elliptical orbit (e < 11) and for $E \ge 0$ the motion is infinite, either on a parabola (E = 0) or a hyperbola (E > 0) with e > 1. The hyperbolic orbit can be expressed via the dependence of r and t on the parameter ξ , the "mean anomaly" of the Kepler orbit, $r = a(e \cosh \xi - 1), \quad t = \sqrt{a^3/Z}(e \sinh \xi - \xi)$ the "semi-axis" of the hyperbola a = Z/2E $e = \sqrt{1 + 2EL^2/Z^2}$



Departmento de Fisica UBACKACCASA The time delay of the classical Kepler trajectory with respect to a freeparticle trajectory can be obtained directly from our asymptotic expression for t(r) $t(r) \stackrel{t}{=} \frac{r}{k} + \Delta t \iff \Delta t(r) \stackrel{z}{=} \frac{Z}{k^3} \left[1 - \ln\left(\frac{2k^2r}{\sqrt{Z^2 + k^2L^2}}\right) \right]$ The time delay as a function of the propagation time is obtained when we express r as a function of t. To first order we take the free particle limit r(t) = kt. $t \frac{t^4_{Coul}(E, L, r = kt) \approx \frac{Z}{k^3} \left[1 - \ln\left(\frac{2k^2t}{\sqrt{\eta^2 + L^2}}\right) \right]$ when we introduce the Coulomb-Sommerfeld parameter $\eta = Z/k$ $t^4_{Coul}(E, L, r = kt) = \frac{Z}{k^3} \ln(\sqrt{\eta^2 + L^2}) + \Delta t_{Coul}(E, r = kt)$ $\Delta t_{Coul}(E, r) = \frac{Z}{k^3} (1 - \ln(2kr))$



Departmento de Fisica UBACKACCESS $t_{Coul}(E, \ell, r) = \frac{\partial}{\partial E} \left(\frac{Z}{k} \ln(2kr) + \sigma_{\ell}(E) \right)$ $= \Delta t_{Coul}(E, r) + t_{EWS}^{C}(E, \ell)$ $\Delta t_{Coul}(E, r) = \frac{Z}{k^3} (1 - \ln(2kr))$ where we use the asymptotic momentum $k = \sqrt{2E}$ $t_{EWS}^{C}(E, \ell) = \frac{\partial}{\partial E} \sigma_{\ell}(E)$ We want to compare the different extraction methods for the EWS time delay using the scattering phase $t_{EWS} = \partial \delta_{\ell}(E)/\partial E$, or the radial wave packet $\langle r(t) \rangle = v_g(t - t_{WP}), \quad t \to \infty$, which are equivalent for short range potentials



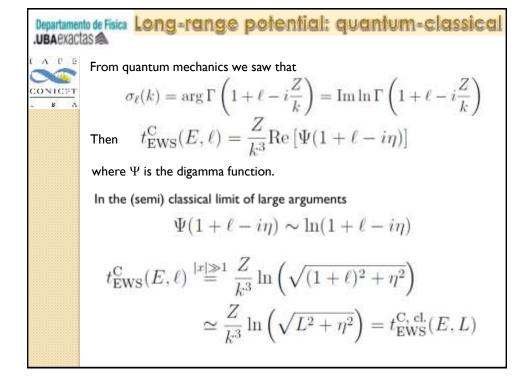
Departmento de Fisica Long-range potential: quantum-classical
UBAEXACTAS
Comparing quantum and classical time delays for long-range potentials:
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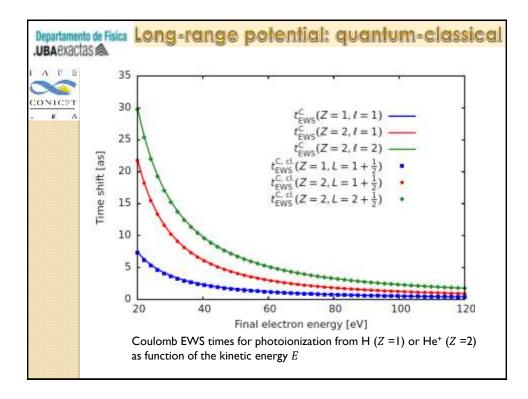
$$t_{Coul}(E, \ell, r) = \frac{\partial}{\partial E} \left(\frac{Z}{k} \ln(2kr) + \sigma_{\ell}(E) \right)$$

$$= \Delta t_{Coul}(E, r) + t_{EWS}^{C}(E, \ell)$$

$$t_{Coul}^{cl}(E, L, r = kt) = \frac{Z}{k^3} \ln(\sqrt{\eta^2 + L^2}) + \Delta t_{Coul}(E, r = kt)$$
suggest to identify the classical analogue of the intrinsic EWS delay

$$t_{EWS}^{C, cl.}(E, L) = \frac{Z}{k^3} \ln\left(\sqrt{\eta^2 + L^2}\right)$$





Departamento de Física Long-range potential: quantum-classical



The Coulomb singularity prohibits electrons at r=0, however, the expectation for an initial state with well-defined parity is still centered around zero (although the radial expectation value $\langle r(t) \rangle \neq 0$). This implies that the time delay is nearly independent of the principal quantum number as long as the initial states have the same angular momentum quantum number.

Since bound states can be chosen to be real, only the phase information of the final state enters in the definition of the EWS delay. As long as only one partial wave contributes, which is the case for initial *s*-states, the EWS time delay is independent of the principal quantum number n.

No matter where the electron initially was, it ends on the same trajectory as long as it has the same final kinetic energy (by choosing appropriate photon energies). The reason is that as long as the initial orbital has a well-defined parity, no matter how large the extension is, the mean position of departure is always at the origin. The radial expectation value $\langle r(t) \rangle$ of the continuum part with E > 0 depends on the main quantum number only as long as the ionization process is not finished (for the duration of the laser pulse).

