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ATTOSECOND PHYSICS

UNIT III


REVIEW OF ELECTROMAGNETISM

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Electromagnetic fields

System of Units: We use MKS despite the most used is the Gaussian

Maxwell Equations

$$\vec{\nabla} \cdot \vec{E} = \frac{\rho}{\epsilon_0}$$

$$\vec{\nabla} \times \vec{E} = -\frac{\partial \vec{B}}{\partial t}$$

$$\vec{\nabla} \cdot \vec{B} = 0$$

$$\vec{\nabla} \times \vec{B} = \mu_0 \vec{J} + \epsilon_0 \mu_0 \frac{\partial \vec{E}}{\partial t}$$

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Electromagnetic fields (cont.)

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$$\vec{\nabla} \cdot \vec{B} = 0 \Rightarrow \boxed{\vec{B} = \vec{\nabla} \times \vec{A}} \quad \vec{A} \text{ is the vector potential}$$

$$\vec{\nabla} \times \vec{E} = -\frac{\partial \vec{B}}{\partial t}$$

$$\vec{\nabla} \times \vec{E} + \frac{\partial \vec{\nabla} \times \vec{A}}{\partial t} = 0$$

$$\vec{\nabla} \times \left(\underbrace{\vec{E} + \frac{\partial \vec{A}}{\partial t}}_{-\vec{\nabla} \phi} \right) = 0 \Rightarrow \boxed{\vec{E} = -\vec{\nabla} \phi - \frac{\partial \vec{A}}{\partial t}}$$

ϕ is the scalar potential

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Gauge Invariance

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Electric and magnetic fields E and B are unaltered if one changes:

$$\boxed{\begin{aligned} \vec{A} &\rightarrow \vec{A}' = \vec{A} + \vec{\nabla} \chi \\ \phi &\rightarrow \phi' = \phi - \frac{\partial \chi}{\partial t} \end{aligned}}$$

where $\chi(\vec{r}, t)$ is an arbitrary function

Proof:

$$\vec{B}' = \vec{\nabla} \times \vec{A}' = \vec{\nabla} \times (\vec{A} + \vec{\nabla} \chi) = \vec{\nabla} \times \vec{A} + \underbrace{\vec{\nabla} \times \vec{\nabla} \chi}_{=0} = \vec{\nabla} \times \vec{A} = \vec{B}$$

$$\vec{E}' = -\vec{\nabla} \phi' - \frac{\partial \vec{A}'}{\partial t} = -\vec{\nabla} \left(\phi - \frac{\partial \chi}{\partial t} \right) - \frac{\partial}{\partial t} (\vec{A} + \vec{\nabla} \chi) = -\vec{\nabla} \phi - \frac{\partial \vec{A}}{\partial t} = \vec{E}$$

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Classical Electromagnetic Radiation Field (cont.)

Solution to the wave equation:

$$\vec{A}_{\lambda\vec{k}} = A_N e^{i(\vec{k}\cdot\vec{r} - \omega t + \delta_\omega)} \hat{\epsilon}_{\lambda\vec{k}} + \text{c.c.}$$

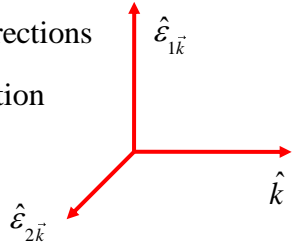
$$= A_N \hat{\epsilon}_{\lambda\vec{k}} 2 \cos(\vec{k}\cdot\vec{r} - \omega t + \delta_\omega)$$

We still have the freedom to restrict :

$$\phi = 0 \Rightarrow \vec{\nabla} \cdot \vec{A} = 0 \quad (\text{Coulomb gauge})$$

$\vec{\nabla} \cdot \vec{A} \propto \vec{k} \cdot \hat{\epsilon}_{\lambda\vec{k}} = 0 \Rightarrow \vec{k} \perp \hat{\epsilon}_{\lambda\vec{k}}$: two directions

When we introduce \vec{A} into the wave equation
 $\Rightarrow \omega = ck$ dispersion relation



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Classical electromagnetic radiation field (cont.)

Electric field: $\vec{E}_{\lambda\vec{k}} = -\frac{\partial \vec{A}_{\lambda\vec{k}}}{\partial t} = A_N \hat{\epsilon}_{\lambda\vec{k}} 2\omega \sin(\vec{k}\cdot\vec{r} - \omega t + \delta_\omega)$

Magnetic field: $\vec{B}_{\lambda\vec{k}} = \vec{\nabla} \times \vec{A}_{\lambda\vec{k}} = A_N (\vec{k} \times \hat{\epsilon}_{\lambda\vec{k}}) 2 \sin(\vec{k}\cdot\vec{r} - \omega t + \delta_\omega)$
 $\Rightarrow \vec{B}_{\lambda\vec{k}} \perp \vec{E}_{\lambda\vec{k}}$ and in phase.

$$\left| \vec{E}_{\lambda\vec{k}} \right| \propto \omega \quad \text{and} \quad \left| \vec{B}_{\lambda\vec{k}} \right| \propto k \quad \Rightarrow \frac{\left| \vec{B}_{\lambda\vec{k}} \right|}{\left| \vec{E}_{\lambda\vec{k}} \right|} = \frac{k}{\omega} = \frac{1}{c}$$

In the non-relativistic regime, as c is big, we can neglect the magnetic field (next slide).

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Classical Electromagnetic Radiation Field (cont.)

The volumetric density of energy

$$\begin{aligned} \rho(\omega) &= \frac{\epsilon_0}{2} \left| \vec{E}_{\lambda k} \right|^2 + \frac{1}{2\mu_0} \left| \vec{B}_{\lambda k} \right|^2 \\ &= \frac{\epsilon_0}{2} A_N^2 4\omega^2 \sin^2 \left(\vec{k} \cdot \vec{r} - \omega t + \delta_\omega \right) + \frac{1}{2\mu_0} A_N^2 4 k^2 \sin^2 \left(\vec{k} \cdot \vec{r} - \omega t + \delta_\omega \right) \\ &= \epsilon_0 A_N^2 4\omega^2 \sin^2 \left(\vec{k} \cdot \vec{r} - \omega t + \delta_\omega \right), \text{ which is time dependent} \\ &\Rightarrow \langle \rho(\omega) \rangle = 2\epsilon_0 A_N^2 \omega^2 \quad (\text{time average}) \end{aligned}$$

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Classical electromagnetic radiation field (cont.)

The Poynting vector

$$\begin{aligned} \vec{S} &= \frac{1}{\mu_0} \vec{E} \times \vec{B} \propto \vec{k} \\ &= \frac{E_0 B_0}{\mu_0} \hat{k} \sin^2 \left(\vec{k} \cdot \vec{r} - \omega t + \delta_\omega \right) \\ &= \frac{E_0^2}{\mu_0 c} \hat{k} \sin^2 \left(\vec{k} \cdot \vec{r} - \omega t + \delta_\omega \right) = \epsilon_0 c E_0^2 \hat{k} \sin^2 \left(\vec{k} \cdot \vec{r} - \omega t + \delta_\omega \right) \end{aligned}$$

Intensity or irradiance: $I = \langle |\vec{S}| \rangle = \frac{\epsilon_0 c E_0^2}{2} \quad [I] = \frac{\text{W}}{\text{m}^2}$

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Classical electromagnetic radiation field (cont.)

Example 1: What is the intensity of the solar radiation on the Earth?

$$I = 1.4 \frac{\text{kW}}{\text{m}^2} = 0.14 \frac{\text{W}}{\text{cm}^2}; \quad E_0 \sim 750 \frac{\text{V}}{\text{m}}; \quad B_0 \sim 2.4 \mu\text{Tesla}$$

Example 2: What is the intensity of the radiation of a He-Ne laser?

Power $\sim 3.2\text{mW}$; diameter beam $\sim 2.5\text{mm}$

$$I \simeq \frac{3.2\text{mW}}{\pi(2.5\text{mm}/2)^2} = 0.065 \frac{\text{W}}{\text{cm}^2}$$

Example 3: What is the intensity of the radiation of a Ti-sapphire laser?

$$I \sim 10^{10} - 10^{16} \frac{\text{W}}{\text{cm}^2}$$

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Classical electromagnetic radiation field (cont.)

Dipole approximation:

Spatial dependence of the vector potential: $\vec{A}_{\lambda k} \propto e^{i\vec{k}\cdot\vec{r}}$ where $k = \frac{\omega}{c} = \frac{2\pi\nu}{c}$

As c is big, we can neglect the wave number k of the photon and, therefore, the spatial dependence of the vector potential
(photon momentum) $\hbar k = \frac{\hbar\omega}{c} = \frac{\varepsilon}{c} \ll 1$

Example: Ti-sapphire laser

$$k = \frac{2\pi}{\lambda} = \frac{2\pi}{8 \times 10^2 \times 10^{-9} \text{m}} = 7.8 \times 10^6 \text{m}^{-1}$$

The dimension of the atom is about some Bohr radii

$$a_0 = 5.3 \times 10^{-11} \text{m}$$

$$ka_0 = \frac{2\pi a_0}{\lambda} = 7.8 \times 10^6 \text{m}^{-1} \times 5.3 \times 10^{-11} \text{m} \simeq 4 \times 10^{-4} \ll 1$$

The dipole approximation fails when:

$$kr \sim 1 \Rightarrow \nu > \frac{c}{2\pi r} \quad \text{with} \quad \frac{c}{2\pi a_0} = 9 \times 10^{17} \text{Hz} \quad (\text{hard X and } \gamma \text{ rays})$$

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Radiation-matter coupling

Lorentz Force: $\vec{F} = q(\vec{E} + \vec{v} \times \vec{B})$

Inserting B and E into the Lorentz law: $\vec{F} = q \left(-\vec{\nabla}\phi - \frac{\partial \vec{A}}{\partial t} + \vec{v} \times \vec{\nabla} \times \vec{A} \right) \quad (1)$

Exercise 4: Prove that (Goldstein pgs. 27-28)

$$\vec{v} \times \vec{\nabla} \times \vec{A} = \vec{\nabla}(\vec{v} \cdot \vec{A}) - \frac{d\vec{A}}{dt} + \frac{\partial \vec{A}}{\partial t} \quad (2)$$

where $\frac{dA_x}{dt} = \frac{\partial A_x}{\partial t} + \underbrace{\left(v_x \frac{\partial A_x}{\partial x} + v_y \frac{\partial A_x}{\partial y} + v_z \frac{\partial A_x}{\partial z} \right)}_{\vec{v} \cdot \vec{\nabla} A_x}$

Replacing (2) in (1) $\vec{F} = q \left(-\vec{\nabla}(\phi - \vec{v} \cdot \vec{A}) - \frac{d\vec{A}}{dt} \right)$

And rewriting $\vec{A} = \vec{\nabla}_v(\vec{v} \cdot \vec{A}) = -\vec{\nabla}_v(\phi - \vec{v} \cdot \vec{A})$

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Lagrangian Formalism

$$\vec{F} = -\vec{\nabla} \left[\underbrace{q(\phi - \vec{v} \cdot \vec{A})}_U \right] + \frac{d}{dt} \left\{ \vec{\nabla}_v \left[\underbrace{q(\phi - \vec{v} \cdot \vec{A})}_U \right] \right\}$$

$$\vec{F} = -\vec{\nabla}U + \frac{d}{dt} \vec{\nabla}_v U \quad \text{Definition of generalized force}$$

$$U = q(\phi - \vec{v} \cdot \vec{A}) \quad U \text{ is the generalized potential}$$

The Lagrangian of a charged particle in an electromagnetic field is given by

$$L(\vec{r}, \dot{\vec{r}}, t) = T_{\text{kin}} - U = \frac{1}{2} m v^2 - q(\phi - \vec{v} \cdot \vec{A}); \quad \text{where } \dot{\vec{r}} = \vec{v}$$

Generalized momentum: $\vec{p} = \frac{\partial L}{\partial \vec{v}}$

$$\vec{p} = m\vec{v} + q\vec{A}; \quad \vec{p} \neq m\vec{v}$$

Watch out!

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Hamiltonian Formalism

By means of a Legendre transform the Hamiltonian is defined by:

$$\begin{aligned}
 H(\vec{p}, \vec{r}, t) &= \dot{\vec{r}} \cdot \vec{p} - L(\vec{r}, \dot{\vec{r}}, t) \\
 &= \vec{v} \cdot \vec{p} - \frac{1}{2}mv^2 + q(\phi - \vec{v} \cdot \vec{A}) \\
 &= \left(\frac{\vec{p} - q\vec{A}}{m} \right) \cdot \vec{p} - \frac{1}{2}m \left(\frac{\vec{p} - q\vec{A}}{m} \right)^2 + q\phi - q \left(\frac{\vec{p} - q\vec{A}}{m} \right) \cdot \vec{A}
 \end{aligned}$$

$$H(\vec{p}, \vec{r}, t) = \frac{(\vec{p} - q\vec{A})^2}{2m} + q\phi$$

Exercise 5: Derive the Lorentz force from the Hamilton equations:

$$\dot{\vec{r}} = \frac{\partial H}{\partial \vec{p}}; \quad -\dot{\vec{p}} = \frac{\partial H}{\partial \vec{r}}$$

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Hamiltonian formalism (cont.)

This is called the length gauge:

$$H'(\vec{p}, \vec{r}, t) = \frac{(\vec{p} - q\vec{A}')^2}{2m} + q\phi' = \frac{\vec{p}^2}{2m} - q\vec{E} \cdot \vec{r}$$

where $\vec{p} = m\vec{v} + q\vec{A}' = m\vec{v}$
=0

In the velocity gauge:

$$H(\vec{p}, \vec{r}, t) = \frac{(\vec{p} - q\vec{A})^2}{2m} + q\phi = \frac{(\vec{p} - q\vec{A})^2}{2m}$$

where $\vec{p} = m\vec{v} + q\vec{A} \neq m\vec{v}$

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Classical electron in a sinusoidal electric field

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CONCEPT

Linearly polarized plane wave (never started and never finished):
 $\vec{E} = \hat{z}E_0 \sin \omega t$; \hat{z} : polarization direction
 $q = -e$; $e = 1.6 \times 10^{-19} \text{ C}$
 Lorentz Force: $m_e \ddot{z} = -eE_0 \sin \omega t$
 $\dot{z} = \frac{eE_0}{m_e \omega} \cos \omega t$ supposing adiabatic switch on and off
 kinetic energy: $T_{\text{kin}} = \frac{1}{2} m_e \dot{z}^2 = \frac{e^2 E_0^2}{2m_e \omega^2} \cos^2 \omega t$
 time average: $\langle T_{\text{kin}} \rangle = \frac{1}{T} \int_t^{t+T} T_{\text{kin}} dt = \frac{e^2 E_0^2}{4m_e \omega^2}$
 $U_p = \frac{e^2 E_0^2}{4m_e \omega^2}$ ponderomotive energy Energy an electron has because it is oscillating driven by E(t)

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Classical electron in a sinusoidal electric field (cont.)

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CONCEPT

Electron position:
 $z(t) = \alpha_0 \sin \omega t$; $\alpha_0 = \frac{eE_0}{m_e \omega^2}$ quiver (oscillation) amplitude
 $\beta = \frac{v}{c} = \frac{eE_0}{m_e \omega c} \Rightarrow \begin{cases} \beta \ll 1 \Rightarrow \text{non-relativistic case} \\ \beta \simeq 1 \Rightarrow \text{relativistic case} \end{cases}$

Exercise 6: How much is β for an electron excited with a Ti-Sapphire laser?
 $\lambda = 800 \text{ nm}$, $E_0 = 5 \times 10^{10} \text{ V/m}$, $m_e = 9.1 \times 10^{-31} \text{ kg}$

Only this century it has been possible to build lasers with $\beta \sim 1$

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Quantum Electron in Electromagnetic Field

CONCEPT

First quantization: $\vec{r} \rightarrow \hat{r}$

$$\vec{p} \rightarrow \hat{p} = \frac{\hbar}{i} \vec{\nabla}$$

$$\varepsilon \rightarrow \hat{H} = i\hbar \frac{\partial}{\partial t}$$

Schrödinger equation: $i\hbar \frac{\partial}{\partial t} \Psi(\vec{r}, t) = \left[\frac{1}{2m_e} \left(\frac{\hbar}{i} \vec{\nabla} - q\vec{A} \right)^2 + q\phi \right] \Psi(\vec{r}, t)$

Exercise 7: How does gauge invariance work for the Schrödinger equation?

$$\vec{A} \rightarrow \vec{A}' = \vec{A} + \vec{\nabla}\chi$$

$$\phi \rightarrow \phi' = \phi - \frac{\partial\chi}{\partial t}$$

Solution: $\Psi'(\vec{r}, t) = \Psi(\vec{r}, t)e^{i\frac{q}{\hbar}\chi}$

In the Coulomb gauge: $\nabla^2\chi = 0$

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Quantum Electron in Electromagnetic Field (cont.)

CONCEPT

As we have already seen $\chi(\vec{r}, t) = -\vec{A} \cdot \vec{r}$

$$\Rightarrow \vec{A}' = 0, \text{ and } \phi' = -\vec{E}(t) \cdot \vec{r}$$

Length gauge: $i\hbar \frac{\partial}{\partial t} \Psi'(\vec{r}, t) = \left[\frac{-\hbar^2}{2m_e} \nabla^2 - q\vec{E}(t) \cdot \vec{r} \right] \Psi'(\vec{r}, t)$

Velocity gauge: $i\hbar \frac{\partial}{\partial t} \Psi(\vec{r}, t) = \frac{1}{2m_e} \left(\frac{\hbar}{i} \vec{\nabla} - q\vec{A} \right)^2 \Psi(\vec{r}, t)$

Let's solve the Schrödinger equation in the velocity gauge.
In the trivial case that there are not electromagnetic fields ($\mathbf{A} = 0$) we have the free particle:

$$i\hbar \frac{\partial}{\partial t} \Psi(\vec{r}, t) = \frac{-\hbar^2}{2m_e} \nabla^2 \Psi(\vec{r}, t)$$

We use the method of separation of variables: $\Psi(\vec{r}, t) = \psi(\vec{r})e^{-\frac{i}{\hbar}\varepsilon_k t}$

$$\frac{-\hbar^2}{2m_e} \nabla^2 \psi(\vec{r}) = \varepsilon_k \psi(\vec{r}) \Rightarrow \psi(\vec{r}) = \frac{e^{i\vec{k} \cdot \vec{r}}}{(2\pi)^{3/2}} \text{ with } \varepsilon_k = \frac{\hbar^2 k^2}{2m_e} \text{ (kinetic energy)}$$

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Quantum Electron in Electromagnetic Field (cont.)

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CONCEPT

Exercise 8: Prove that if $\vec{A} = \vec{A}(t) \neq 0 \Rightarrow \Psi(\vec{r}, t) = \psi(\vec{r}) e^{-\frac{i}{\hbar} S(t)}$
where $S(t) = \frac{1}{2m_e} \int_{t_0}^t dt' \left[\hbar \vec{k} - q \vec{A}(t') \right]^2$ classical action

Let's consider the case of a linearly polarized sinusoidal electric field with adiabatic switch on and off

$$\vec{E}(t) = \hat{z} E_0 \sin \omega t$$

$$\vec{E}(t) = -\frac{\partial \vec{A}}{\partial t} \Rightarrow \vec{A}(t) = -\int_{-\infty}^t dt' \vec{E}(t') = -\hat{z} E_0 \int_{-\infty}^t dt' \sin \omega t' = \hat{z} \frac{E_0}{\omega} \cos \omega t$$

For the sinusoidal case we can derive an analytical action:

$$S(t) = \frac{1}{2m_e} \int_{t_0}^t dt' \left[\hbar \vec{k} - q \vec{A}(t') \right]^2 = \varepsilon_k(t - t_0) + \hbar \vec{k} \cdot \vec{\alpha}(t) + \beta(t)$$

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Quantum Electron in Electromagnetic Field (cont.)

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CONCEPT

$$\vec{\alpha}(t) = \frac{-q}{m_e} \int_{t_0}^t dt' \vec{A}(t') = \frac{-q}{m_e} \frac{E_0}{\omega} \hat{z} \int_{t_0}^t dt' \cos \omega t' = \frac{-q}{m_e} \frac{E_0}{\omega^2} \hat{z} \sin \omega t \quad (\text{adiabatic switch on})$$

is the integral of the integral of the force \Rightarrow displacement of the classical electron

$$\langle \vec{\alpha}(t) \rangle = \frac{1}{T} \int_0^T dt \vec{\alpha}(t) = 0$$

$$\beta(t) = \frac{q^2}{2m_e} \int_{t_0}^t dt' A^2(t') = \frac{q^2}{2m_e} \frac{E_0^2}{\omega^2} \int_{t_0}^t dt' \cos^2 \omega t' = \frac{q^2}{4m_e} \frac{E_0^2}{\omega^2} \left[(t - t_0) + \frac{\sin 2\omega t}{2\omega} \right]$$

where I used adiabatic switch on

$$\langle \beta(t) \rangle = \frac{1}{T} \int_0^T dt \beta(t) = \frac{q^2}{4m_e} \frac{E_0^2}{\omega^2} = U_p \quad (\text{ponderomotive energy})$$

$q = -e$ for the electron

If $E_0 = 0$ everything reduces to the free particle