

31/08 Relatividad General: Práctica

Clase 1: Guía 1 Relatividad Especial

- Convenciones de la práctica:

- Indices griegos indican coordenadas espacio-temporales

$$\mu, \nu, \rho, \sigma = 0, 1, 2, 3$$

↓
tiempo espacio

ejemplo: $x^{\mu=0} = x^0 = ct$, $x^2 = y$

- Indices latinos indican coordenadas espaciales

$$i, j, k = 1, 2, 3$$

ejemplo: $x^{j=3} = x^3 = z$

- Signatura de la métrica: mostly plus

Minkowski

$$\eta_{\mu\nu} = \begin{bmatrix} \boxed{-1} & & & \\ & +1 & & \\ & & +1 & \\ & & & +1 \end{bmatrix}$$

temporal

- unidades $\boxed{c=1} \Rightarrow$ tiempo y longitudes tienen las mismas unidades

Cuadrivectores

$$\eta_{\mu\nu} = \begin{bmatrix} -1 & & & \\ & +1 & & \\ & & +1 & \\ & & & +1 \end{bmatrix}$$

V : Cuadrivector

$$V \cdot V \equiv V_{\mu} V^{\mu} = \eta_{\mu\nu} V^{\mu} V^{\nu} \quad \text{norma de } V$$

convención suma de Einstein $V_{\mu} V^{\mu} = \sum_{\mu=0}^3 V_{\mu} V^{\mu}$

$\eta_{\mu\nu}$ provee un producto escalar para vectores q' NO es definido positivo

$$V \cdot V > 0$$

V es tipo espacio (spacelike)

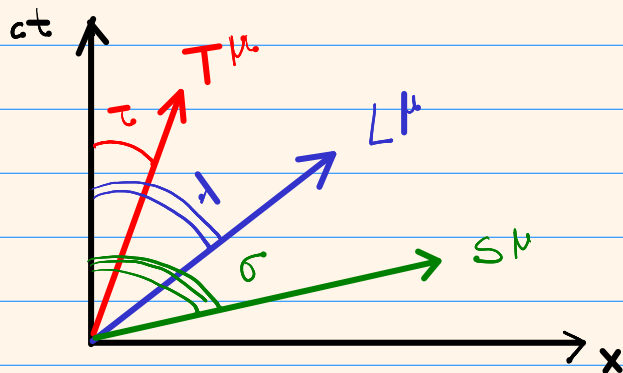
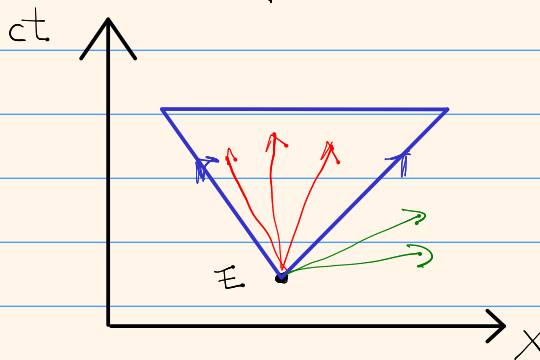
$$V \cdot V = 0$$

V es tipo luz (lightlike or null)

$$V \cdot V < 0$$

V es tipo tiempo (timelike)

• Diagrama espacio-temporal

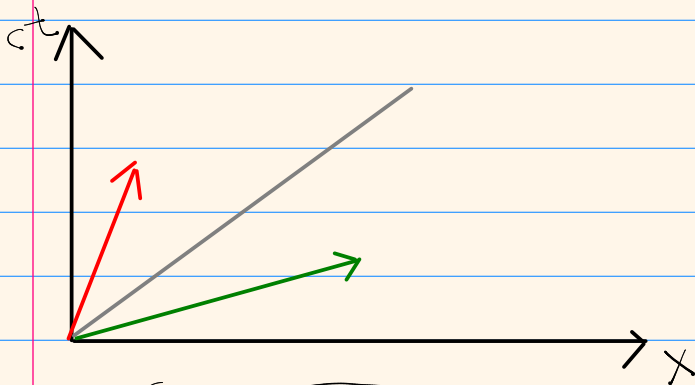


$$\tau < \pi/4 \quad T_{\mu} T^{\mu} < 0$$

$$\lambda = \pi/4 \quad L_{\mu} L^{\mu} = 0$$

$$\sigma > \pi/4 \quad S_{\mu} S^{\mu} > 0$$

$$\tan \alpha = \frac{\text{"comp } x}{\text{"comp } t"}$$



2D

$$\|T\|^2 = -(T^0)^2 + (T^x)^2 < 0$$

$$T^0 = \|T\| \operatorname{ch}(\Theta_T)$$

$$T^x = \|T\| \operatorname{sh}(\Theta_T)$$

$$\|S\|^2 = -(S^0)^2 + (S^x)^2 > 0$$

$$S^0 = \|S\| \operatorname{sh}(\Theta_S)$$

$$S^x = \|S\| \operatorname{ch}(\Theta_S)$$

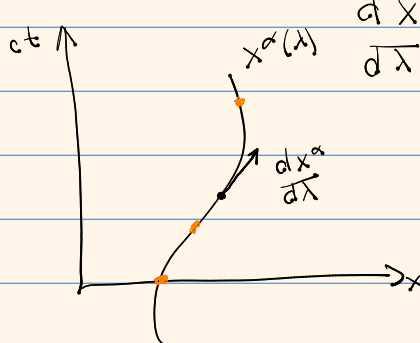
(ver Ferraro sec 7.3)

Resumen:

	tipo de signatura vector	tiempo	nulo	espacio
práctica	mostly plus	$T \cdot T < 0$	$L \cdot L = 0$	$S \cdot S > 0$
teórica	mostly minus	$T \cdot T > 0$	$L \cdot L \leq 0$	$S \cdot S < 0$

Problema 1

trayectoria de un detector $x^\alpha(\lambda)$
 vector tangente a la trayectoria



$$\frac{dx^\alpha(\lambda)}{d\lambda} = \gamma(v) \begin{pmatrix} c \cosh\left(\frac{a\lambda}{c}\right) \\ c \sinh\left(\frac{a\lambda}{c}\right) \\ 0 \\ v \end{pmatrix}$$

$\gamma(v) \equiv (1 - v^2/c^2)^{-1/2}$

(a partir de ahora tomo $c=1$)

a) ver q! λ es el tiempo propio

tiempo propio $\left(d\tau^2 \right) \equiv -ds^2 = -\eta_{\mu\nu} dx^\mu dx^\nu$

intervalo

τ : tiempo propio parametriza curvas temporales.

escribimos el intervalo

$$ds^2 = \eta_{\mu\nu} dx^\mu dx^\nu = \eta_{\mu\nu} \frac{dx^\mu}{d\lambda} \frac{dx^\nu}{d\lambda} d\lambda^2 = \left(-\left(\frac{dx^0}{d\lambda}\right)^2 + \left(\frac{dx^1}{d\lambda}\right)^2 + \left(\frac{dx^2}{d\lambda}\right)^2 + \left(\frac{dx^3}{d\lambda}\right)^2 \right) d\lambda^2$$

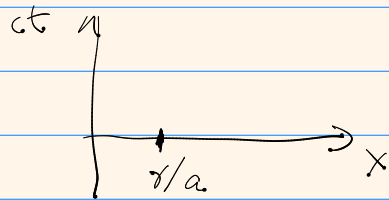
$$= \gamma^2(v) \left(-\cosh^2(a\lambda) + \sinh^2(a\lambda) + v^2 \right) d\lambda^2 = \gamma^2(v) \underbrace{(-1 + v^2)}_{= -\gamma^{-2}(v)} d\lambda^2$$

$$ds^2 = -d\lambda^2 = -d\tau^2 \Rightarrow \boxed{\lambda = \tau}$$

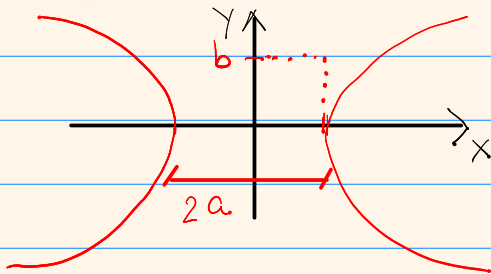
b) hallar $x^\alpha(\tau)$ $\rightarrow$ simplemente integrales ($v = \text{const.}$)

$$x^\alpha(\tau) = \int \frac{dx^\alpha}{d\tau} d\tau$$

$$\begin{cases} x^0(\tau) = \frac{\gamma(v)}{a} \text{sh}(a\tau) + \cancel{\omega\tau} \\ x^1(\tau) = \gamma/a \text{ch}(a\tau) + \cancel{\omega\tau} \\ x^2(\tau) = \cancel{\omega\tau} \\ x^3(\tau) = v\tau + \cancel{\omega\tau} \end{cases}$$



Hipérbola

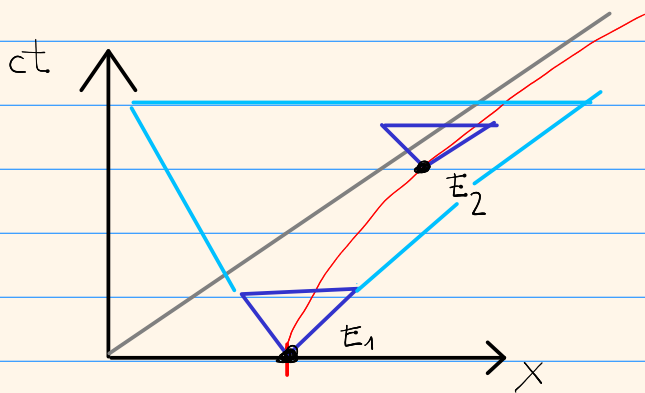


$$\frac{x^2}{a^2} - \frac{y^2}{b^2} = 1$$

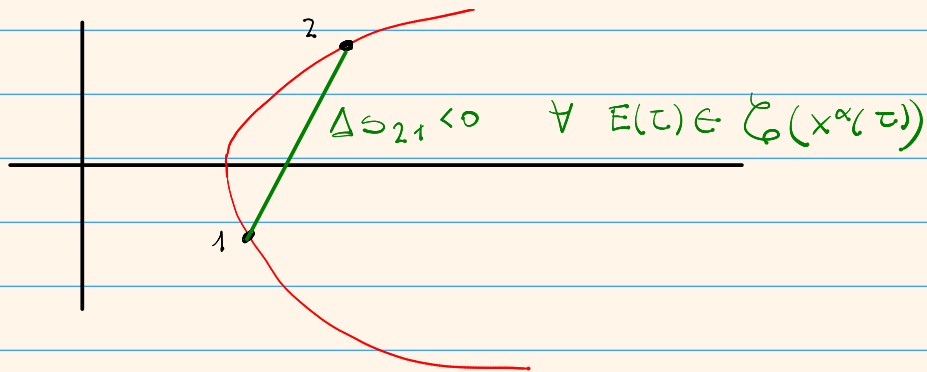
$$-(x^0)^2 + (x^1)^2 = -\frac{\gamma^2}{a^2} \text{sh}^2(a\tau) + \frac{\gamma^2}{a^2} \text{ch}^2(a\tau) = \frac{\gamma^2}{a^2} > 0$$

$$\boxed{-(x^0(\tau))^2 + (x^1(\tau))^2 = \gamma^2/a^2 > 0} \quad \text{hipérbola en } (t, x)$$

c) $x^\alpha(\tau)$ describe una curva temporal?



Para corroborar q: $x^\alpha(\tau)$ describe una curva tipo tiempo la separación entre dos eventos cualesquiera pertenecientes a trayectoria tiene q: ser negativo



la separación entre eventos es temporal

$$\Delta s_{21}^2 = \|x^\mu(\tau_2) - x^\mu(\tau_1)\|_{\text{Mink}}^2 = \eta_{\mu\nu} (x^\mu(\tau_2) - x^\mu(\tau_1)) (x^\nu(\tau_2) - x^\nu(\tau_1)) < 0$$

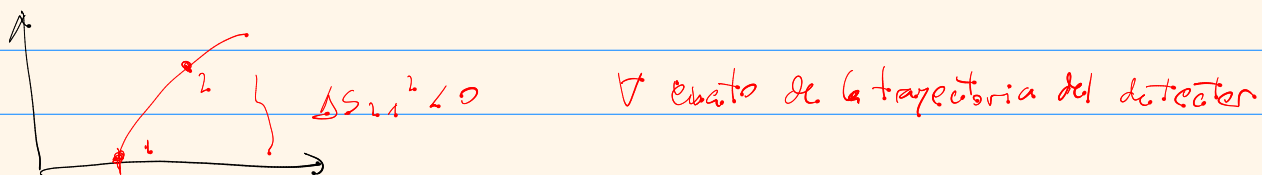
$$= - (x^0(\tau_2) - x^0(\tau_1))^2 + (x^1(\tau_2) - x^1(\tau_1))^2 + (x^3(\tau_2) - x^3(\tau_1))^2$$

$$= -\gamma^2(v) (sh(a\tau_2) - sh(a\tau_1))^2 + \gamma^2(v) (ch(a\tau_2) - ch(a\tau_1))^2 + \gamma^2(v) v^2 \overbrace{(\tau_2 - \tau_1)^2}^{\Delta\tau}$$

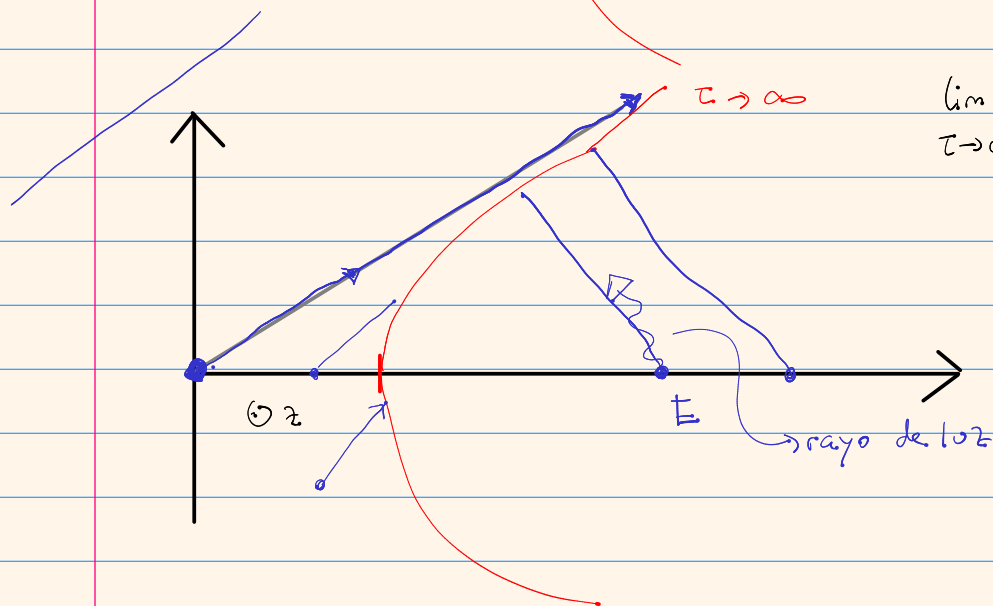
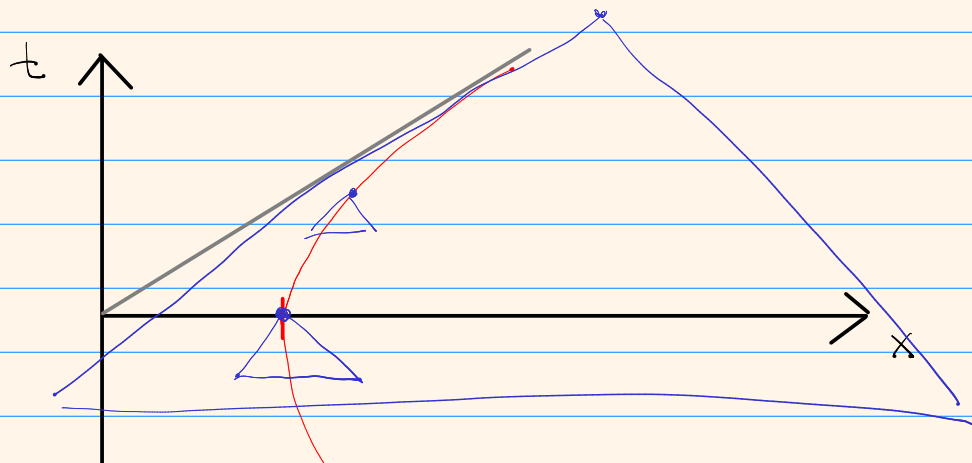
$$= \gamma^2(v) \left\{ -sh^2(a\tau_2) - sh^2(a\tau_1) + 2sh(a\tau_2)sh(a\tau_1) + ch^2(a\tau_2) + ch^2(a\tau_1) - 2ch(a\tau_2)ch(a\tau_1) + v^2 \Delta\tau^2 \right\}$$

$$\Delta s_{21}^2 = \gamma^2(v) \left\{ \underbrace{2 - 2ch(a\Delta\tau)}_{< 0} \right\} + \gamma^2(v) v^2 \Delta\tau^2 < 0$$

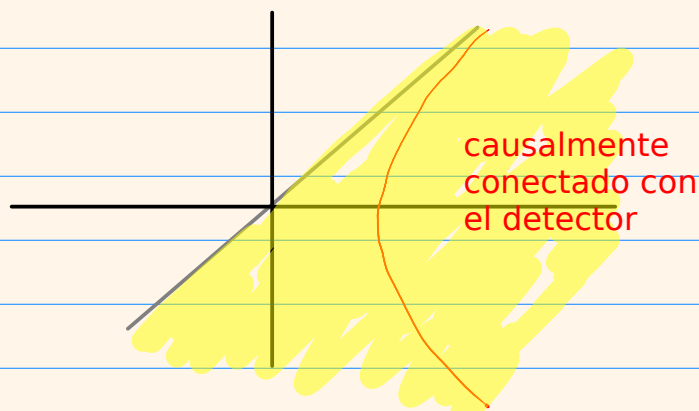
$ch(ax) \cong 1 + \frac{a^2 x^2}{2}$



c) Región del espacio tiempo q. puede influir s/ el detector



$$\lim_{\tau \rightarrow \infty} \frac{x^0(\tau)}{x^1(\tau)} = 1 \text{ (recta a } 45^\circ)$$



e) Temperatura

$$T(\underline{t}, x, y, z) = \frac{Q' t}{x^2 + y^2}$$

calcular $\frac{dT}{d\tau}$

↳ velocidad $U^\mu \equiv \frac{dx^\mu}{d\tau}$

$$\downarrow \quad \downarrow$$
$$(x^1(\tau))^2 \quad (x^2(\tau))^2$$

$$\frac{dT}{d\tau} = \left(\frac{dx^\mu}{d\tau} \right) \frac{\partial T}{\partial x^\mu}(t, x, y, z) = U^\mu \partial_\mu T$$

$$\partial_\mu = \frac{\partial}{\partial x^\mu}$$

$$= U^0 \partial_t T + U^x \partial_x T + U^y \partial_y T + U^z \partial_z T$$

$$\downarrow \quad \downarrow$$
$$\frac{dx^\alpha}{d\tau}$$

$$\frac{\partial}{\partial x^0} = \frac{\partial}{\partial t}$$

Comentarios s/ problemas 2, 3 y 4

Notación: transf de Lorentz $\underbrace{ct'}_{x^{0'}} = \gamma(V) \underbrace{(ct - \beta x)}_{x^0}, x' = \gamma(V) \underbrace{(x - \beta ct)}_{x^1}$

se suele denotar $x^{0'} = \gamma(V)(x^0 - \beta x^1), x^{1'} = \gamma(V)(x^1 - \beta x^0)$

$$\Rightarrow \overset{(S)}{\curvearrowright} x^{\mu'} = \Lambda^{\mu'}_{\nu} \overset{(S)}{\curvearrowright} x^{\nu} \quad x^{\mu'} \equiv (x')^{\mu'}$$

Coords en S x^{μ} , Coords en S' $x^{\mu'}$

Ejemplo: $T_{\mu\nu} \Rightarrow T'_{\mu'\nu'} = \Lambda^{\mu}_{\mu'} \Lambda^{\nu}_{\nu'} T_{\mu\nu}$

$$U^{\mu} = \frac{dx^{\mu}}{d\tau}$$

$$a^{\mu} = \frac{dU^{\mu}}{d\tau} = \frac{d^2 x^{\mu}}{d\tau^2}$$

U^{μ} 4-vectores U^0, U^1, U^2

$$\vec{U}^{\mu} = (U^0, \vec{U}^i) \rightsquigarrow \vec{u}$$

$u^{\mu} = \frac{dx^{\mu}}{dt} = (\dot{x}^0, \vec{u})$ 3-vector $\uparrow$

$$a^{\mu} = \frac{du^{\mu}}{dt} = \frac{d^2 x^{\mu}}{dt^2}$$

$$U^{\mu} = \frac{dx^{\mu}}{d\tau} = \frac{dt}{d\tau} \frac{dx^{\mu}}{dt} = \underbrace{\left(\frac{dt}{d\tau}\right)}_{\gamma} u^{\mu} = \gamma(u) u^{\mu}$$