

Ensembles

Lectura: M. Kardar Cap. 4; R. K. Pathria & D. Beale Caps. 3.

» Ensamble canónico

Cuando el sistema (S) no se encuentra aislado y puede intercambiar calor con un reservorio \mathcal{R} , la energía del sistema no es constante.

Pero $t = \mathcal{R} + S$ está cerrado, y además con $E_t \gg E_S$, la probabilidad de un microestado μ_t en $\mathcal{R} + S$ es

$$P(\mu_S \otimes \mu_{\mathcal{R}}) = \frac{1}{\Omega_{S+\mathcal{R}}(E_t)} \Rightarrow P(\mu_S) = \sum_{\mu_{\mathcal{R}}} P(\mu_S \otimes \mu_{\mathcal{R}}) = \frac{\Omega_{\mathcal{R}}(E_t - E_S)}{\Omega_{S+\mathcal{R}}}$$

O sea, $P(\mu_S) \propto e^{\frac{1}{k_B} S_{\mathcal{R}}(E_t - E_S)}$. Desarrollo $S_{\mathcal{R}}$ como

$$S_{\mathcal{R}}(E_t - E_S) \simeq S_{\mathcal{R}}(E_t) - E_S \frac{\partial S_{\mathcal{R}}}{\partial E_{\mathcal{R}}} = S_{\mathcal{R}}(E_t) - \frac{E_S}{T}$$

$$P(\mu_S) = \frac{e^{-\beta E_S}}{Z(T)}, \quad \text{con} \quad Z(T) = \sum_{\mu_S} e^{-\beta E_S}$$

Función de Partición Canónica

Energía Interna U

$$U = \langle H \rangle = \sum_s E_s \frac{e^{-\beta E_s}}{Z}$$
$$= -\frac{1}{Z} \frac{\partial}{\partial \beta} \sum_s e^{-\beta E_s}$$
$$U = -\frac{\partial}{\partial \beta} \ln Z$$

La energía libre de Helmholtz

$$F = U - TS \Rightarrow U = F + TS$$

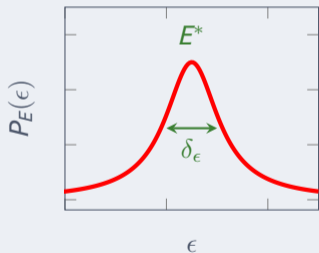
$$= F - T \frac{\partial F}{\partial T} \Big|_{N,V} = -T^2 \frac{\partial}{\partial T} \left(\frac{F}{T} \right) \Big|_{N,V}$$

$$U = \frac{\partial(\beta F)}{\partial \beta} \Big|_{N,V}$$

$-\beta F = \ln Z(T, N, \dots)$ Toda la Termodinámica

» La distribución de probabilidad de las energías $P_E(\epsilon)$

$$P_E(\epsilon) = \sum_s P(\mu_s) \delta(E_s - \epsilon) = \frac{e^{-\beta\epsilon}}{Z} \Omega(\epsilon) = \frac{1}{Z} e^{\left[\frac{S(\epsilon)}{k_B} - \frac{\epsilon}{k_B T} \right]} = \frac{1}{Z} e^{-\beta \tilde{F}(\epsilon)}$$



con $\tilde{F}(\epsilon) = \epsilon - TS(\epsilon)$, ¿Quién es E^* ?

$$\left. \frac{\partial \tilde{F}}{\partial \epsilon} \right|_T = 1 - T \left. \frac{\partial S}{\partial \epsilon} \right|_T \Rightarrow$$

$$\frac{1}{T} = \left. \frac{\partial S(\epsilon)}{\partial \epsilon} \right|_{T, \epsilon=E^*} \Rightarrow$$

$$E^* \stackrel{?}{=} U = \int d\epsilon P_E(\epsilon) \epsilon \quad \text{en el límite termodinámico??}$$

¿Qué pasa con δ_ϵ^2 , i.e., $\langle H^2 \rangle - \langle H \rangle^2$

$$\langle H^2 \rangle - \langle H \rangle^2 = \int \epsilon^2 P_E(\epsilon) d\epsilon - \left(\int \epsilon P_E(\epsilon) d\epsilon \right)^2 = \int \epsilon^2 e^{-\beta\epsilon} \frac{\Omega(\epsilon)}{Z} - \left[\int \epsilon e^{-\beta\epsilon} \frac{\Omega(\epsilon)}{Z} \right]^2$$

pero, $Z(\beta) = \int e^{-\beta\epsilon} \Omega(\epsilon) d\epsilon$

$$\begin{aligned} \frac{\partial^2 \ln Z}{\partial \beta^2} &= -\frac{\partial U}{\partial \beta} = -\frac{\partial}{\partial \beta} \left[\int \epsilon e^{-\beta\epsilon} \frac{\Omega(\epsilon)}{Z} d\epsilon \right] = \int \epsilon^2 e^{-\beta\epsilon} \frac{\Omega(\epsilon)}{Z} d\epsilon + \int \epsilon e^{-\beta\epsilon} \Omega(\epsilon) \frac{\partial Z}{\partial \beta} \frac{1}{Z^2} \\ &= \langle H^2 \rangle + \frac{\partial \ln Z}{\partial \beta} \int \epsilon e^{-\beta\epsilon} \Omega(\epsilon) \frac{1}{Z} = \langle H^2 \rangle - \langle H \rangle^2 \end{aligned}$$

Pero además,

$$-\frac{\partial U}{\partial \beta} = -\frac{\partial U}{\partial T} \frac{\partial T}{\partial \beta} = k_B T^2 \frac{\partial U}{\partial T} = k_B T^2 C \quad \text{es extensivo, i.e., } \propto N$$

o sea

$$\boxed{\frac{\langle H^2 \rangle - \langle H \rangle^2}{U^2} \propto \frac{1}{N}}$$

» y la $P_E(\epsilon)$, entonces?

$$P_E(\epsilon) = \frac{1}{Z} e^{\beta(S(\epsilon)T - \epsilon)} = e^{-\beta\tilde{F}(\epsilon)}$$

Desarrollando \tilde{F} alrededor de $E^* = U$, para aproximar $P_E(\epsilon)$ por una gaussiana,

$$\begin{aligned}\tilde{F}(\epsilon) &= F(U) + \left. \frac{\partial \tilde{F}}{\partial \epsilon} \right|_{\epsilon=U} (\epsilon - U) + \frac{1}{2} \left. \frac{\partial^2 \tilde{F}}{\partial \epsilon^2} \right|_{\epsilon=U} (\epsilon - U)^2 \\ &= F(U) + 0 - \frac{1}{2} T \left. \frac{\partial^2 S}{\partial \epsilon^2} \right|_{\epsilon=U} (\epsilon - U)^2 \\ &= F(U) + \frac{1}{2T} \frac{\partial T}{\partial U} (\epsilon - U)^2 \\ &= F + \frac{1}{2TC} (\epsilon - U)^2\end{aligned}$$

O sea,

$$P_E(\epsilon) \simeq \frac{1}{Z} e^{-\beta F} e^{-\frac{\beta}{2TC} (\epsilon - U)^2}$$

» Hamiltoniano Clásico

Sea un Hamiltoniano clásico $H = H(\{p, q\})$, la suma sobre microestados va con

$$d\Gamma_N = \frac{1}{N!} \prod_{i=1}^N \frac{d^3 q_i d^3 p_i}{h^3}, \quad \text{con lo cual}$$

$$Z(T) = \int d\Gamma_N e^{-\beta H(\{p, q\})}, \quad U = -\frac{\partial \ln Z}{\partial \beta}, \quad F = -k_B T \ln Z$$

Hamiltonianos separables

Si $H = \sum H_k$

$$Z(T) = \frac{1}{N!} \prod_k \left(\int d\Gamma_N^{(k)} e^{-\beta H_k} \right)$$

$$F = \sum_k F_k \quad \& \quad U = \sum_k U_k$$

- Sistemas no interactuantes.
- Partículas libres.
- Grados de libertad internos.

» Gas Ideal

Para un Hamiltoniano de N partículas libres no interactuantes, en $\{p_i, q_i\}$,

$$H = \sum_{i=1}^N \frac{p_i^2}{2m}.$$

$$\begin{aligned} Z(T, V, N) &= \frac{1}{N!} \int \prod_{i=1}^N \frac{d^3 p_i d^3 q_i}{h^3} e^{-\beta \sum \frac{p_i^2}{2m}} = \frac{1}{N!} \prod_{i=1}^N \int \frac{d^3 p_i d^3 q_i}{h^3} e^{-\beta \frac{p_i^2}{2m}} \\ &= \frac{V^N}{N!} \left[\int \frac{d^3 p}{h^3} e^{-\beta \frac{p^2}{2m}} \right]^N = \frac{V^N}{N!} \frac{1}{\lambda(T)^{3N}} \simeq \left(\frac{Ve}{N\lambda^3} \right)^N \Rightarrow \end{aligned}$$

$$F(T, V, N) = -k_B T \ln Z = -k_B N T \left\{ \ln \left[\frac{V}{N\lambda^3} \right] + 1 \right\}$$

$$U = -\frac{\partial \ln Z}{\partial \beta}$$

» Espines en un campo magnético

N espines localizados, no interactuantes, que pueden orientarse

arbitrariamente, $E = \sum_{i=1}^N E_i = \sum_i \mathbf{m}_i \cdot \mathbf{H} = -\mu_0 H \sum_i \cos \theta_i$. La función de partición es

$$Z_N(\beta) = [Z_1(\beta)]^N$$

$$Z_1(\beta) = \int_0^{2\pi} d\phi \int_0^\pi e^{\beta\mu_0 H \cos \theta} \sin \theta d\theta = 2\pi \int_{-1}^1 e^{\beta\mu_0 H x} dx = \frac{2\pi}{\beta\mu_0 H} \left(e^{\beta\mu_0 H} - e^{-\beta\mu_0 H} \right)$$

$$= \frac{4\pi}{\beta\mu_0 H} \sinh \beta\mu_0 H$$

¿y la magnetización media $M = N \langle \mu_0 \cos \theta \rangle$?

$$M_z = \frac{N}{\beta} \frac{\partial}{\partial H} \ln Z_1 = - \frac{\partial F}{\partial H} \Big|_T$$

» Espines en un campo magnético (Cont.)

$$\frac{M}{N} = \langle \mu_0 \cos \theta \rangle = \mu_0 \left[\coth(\beta \mu_0 H) - \frac{1}{\beta \mu_0 H} \right]$$

Si $\beta \mu_0 H \ll 1$ (altas T),

$$M = N \frac{\mu_0^2}{3} \beta H$$

$$\chi_T = \lim_{H \rightarrow 0} \frac{\partial M}{\partial H} \Big|_T \simeq N \frac{\mu^2}{3k_B T} = \frac{\text{Cte}}{T} \quad \text{Ley de Curie}$$

